

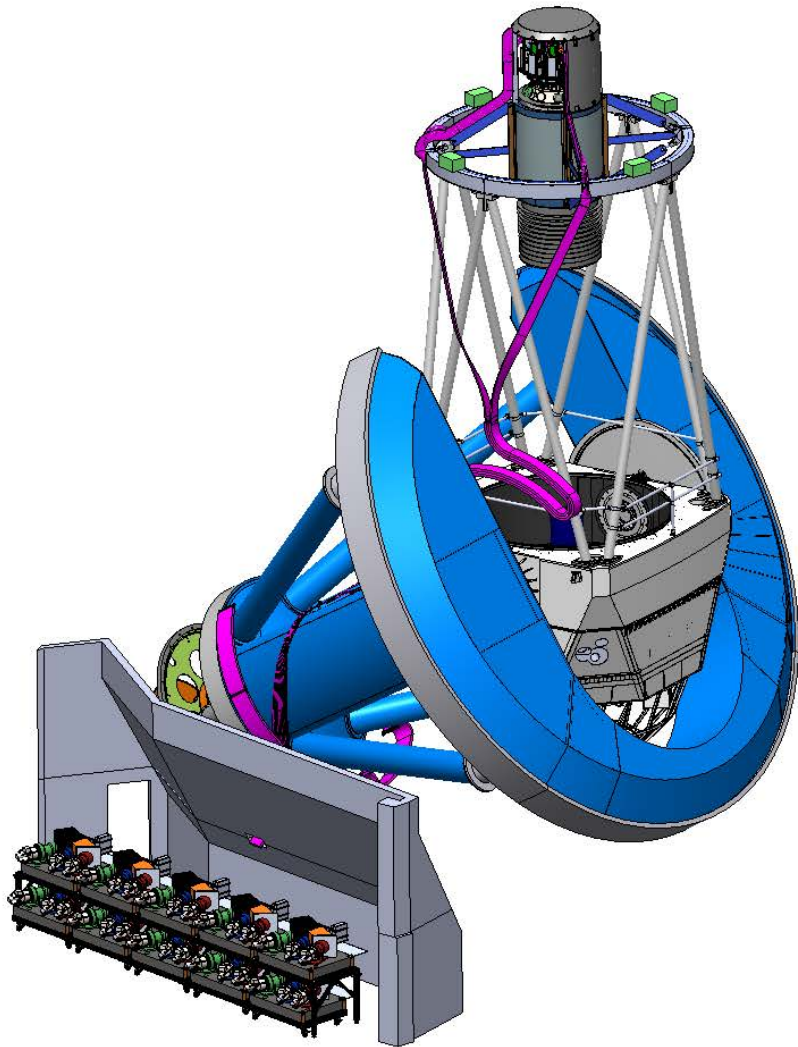
The Dark Energy Spectroscopic Instrument, DESI in 10mins

Will Percival
April 9, 2020

<https://www.desi.lbl.gov/>

<https://arxiv.org/abs/1611.00036>

<https://arxiv.org/abs/1611.00037>



Dark Energy Spectroscopic Instrument

U.S. Department of Energy Office of Science
Lawrence Berkeley National Laboratory



DESI contributors (<https://www.desi.lbl.gov/desi-builders>)

Aguilar	Jessica	leadership of petal fabrication and assembly	Guy	Julien	leadership as Project Scientist and contributions	Poppett	Claire	leadership of the fiber system
Ameel	Jon	contributions to the focal plane electronics	to the data pipeline			Prada	Francisco	contributions to the fiber positioners
Bailey	Stephen	leadership of the data systems and data	Harris	Stewart	leadership as Systems Engineer	Probst	Ronald	contributions to the telescope and installation work
pipeline development			Heetderks	Ian	contributions to the construction of the focal plane	Rabinowitz	David	contributions to the fiber view camera
Baltay	Charles	contributions to the fiber view camera	Heetderks	Henry	leadership as Project Manager	Raichoor	Anand	contributions to the target selection
Bebek	Chris	leadership as Instrument Scientist	Herrera	David	contributions to the imaging survey	Rockosi	Constance	leadership as Commissioning Scientist
Besuner	Robert	leadership as Project Manager and Systems Engineer	Honscheid	Klaus	leadership of the development of the instrument control system	Roe	Natalie	contributions to the project management
Blanc	Pierre-Eric	contributions to the spectrographs	Ishikawa	Yuzo	contributions to the spectrographs	Ronayette	Samuel	contributions to the spectrographs
Blum	Robert	contributions to project management and governance	Jelinsky	Patrick	leadership of the construction of the spectrographs	Ross	Ashley	contributions to the commissioning instrument
Bramall	David	contributions to the fiber system	Jelinsky	Sharon	contributions to the construction of the focal plane	Schlafly	Edward	contributions to the imaging survey and to the focal plane
Brooks	David	contributions of the wide-field corrector control system	Jimenez	Jorge	contributions to the construction of the guiders	Schlegel	David	leadership as Project Scientist and contributions to the imaging survey
Buckley-Geer	Elizabeth	contributions to the instrument control system	Joyce	Richard	leadership as Instrument Scientist for	Schubnell	Michael	contributions to the fiber positioners
Burleigh	Kaylan	contributions to the imaging survey	NOAO/Mayall			Sharples	Ray	leadership of the fiber system
Cahn	Robert	contributions to the design of the survey	Karcher	Armin	contributions to the detector electronics	Silber	Joseph	leadership of the construction of the focal plane
Carton	Pierre-Henri	contributions to the spectrographs and cryostats	Karkar	Sonia	contributions to the spectrographs	Sprayberry	David	leadership of the installation of the corrector and instrument
Claybaugh	Todd	contributions to the focal plane system	Kehoe	Robert	contributions to the data system and to project management	Staten	Ryan	contributions to the data system
Cooper	Andrew	contributions to the support of the bright galaxy and milky way science	Kent	Stephen	contributions to the focal plane software	Tarle	Gregory	contributions to the fiber positioners and for leadership as institutional board chair
Dawson	Kyle	contributions to the design of the survey and survey validation	Kirkby	David	contributions to the survey design	Valdes	Francisco	contributions to the imaging survey
Derwent	Mark	contributions to the spectrographs	Kisner	Theodore	contributions to the pipeline software	Weaver	Benjamin	contributions to the software infrastructure
Dey	Arjun	leadership of the observations and contributions to the imaging surveys	Kitanidis	Ellie	contributions to the imaging survey	Weaverdyck	Curtis	contributions to the fiber positioners
Dhungana	Govinda	contributions to the data system	Lambert	Andrew	contributions to the focal plane system	Wechsler	Risa	leadership as collaboration Spokesperson and contributions to survey leadership
Dobson	Carl	contributions to the instrument software	Landriau	Martin	contributions to the imaging survey	Yeche	Christophe	contributions to the target selection and survey validation
Doel	Peter	leadership of the construction of the wide-field corrector	Lang	Dustin	contributions to the imaging survey	Zhang	Kai	contributions to the focal plane
Duan	Yutong	contributions to the focal plane system	Lee	Jae Hyeon	contributions to target selection	Zhang	Tianmeng	contributions to the imaging survey
Edelstein	Jerry	contributions to the spectrographs	Levi	Michael	leadership of the DESI project as Director	Zhou	Zhimin	contributions to the BASS imaging survey
Eftekharzadeh	Sarah	contributions to the survey design	Magneville	Christophe	contributions to the spectrographs and cryostats	Zhou	Xu	contributions to the BASS imaging survey
Eisenstein	Daniel	leadership as collaboration Spokesperson and contributions to the design of the survey	Martini	Paul	leadership as instrument scientist and contributions to the commissioning instrument	Zhou	Rongpu	contributions to the target selection
Elliott	Ann	contributions to the instrument operation	McDonald	Patrick	contributions to the survey design	Zou	Hu	contributions to the BASS imaging survey
pipeline			Meisner	Aaron	contributions to the imaging survey			
Estrada	Juan	contributions to the CCD packaging	Miller	Tim	contributions to the wide-field corrector			
Fagrellius	Parker	contributions to the instrument and its commissioning	Moustakas	John	contributions to the imaging survey			
Fanning	Kevin	contributions to the focal plane system	Myers	Adam	leadership of the target selection pipeline			
Flaugher	Brenna	leadership as Project Scientist and leadership of the AT&C subsystem	Newman	Jeffrey	contributions to the target selection			
Gaztanaga	Enrique	contributions to the imaging validation	Nie	Jundan	contributions to the imaging survey			
Gutierrez	Gaston	leadership of the construction of the wide-field corrector	Palanque-Delabrouille	Nathalie	leadership as Spokesperson and contributions to target selection			
			Pappalardo	Daniel	contributions to the spectrographs			
			Peng	Xiyun	contributions to the imaging survey			
			Percival	Will	contributions to the preparation of the galaxy and quasar clustering science			



Dark Energy Spectroscopic Instrument

U.S. Department of Energy Office of Science
Lawrence Berkeley National Laboratory

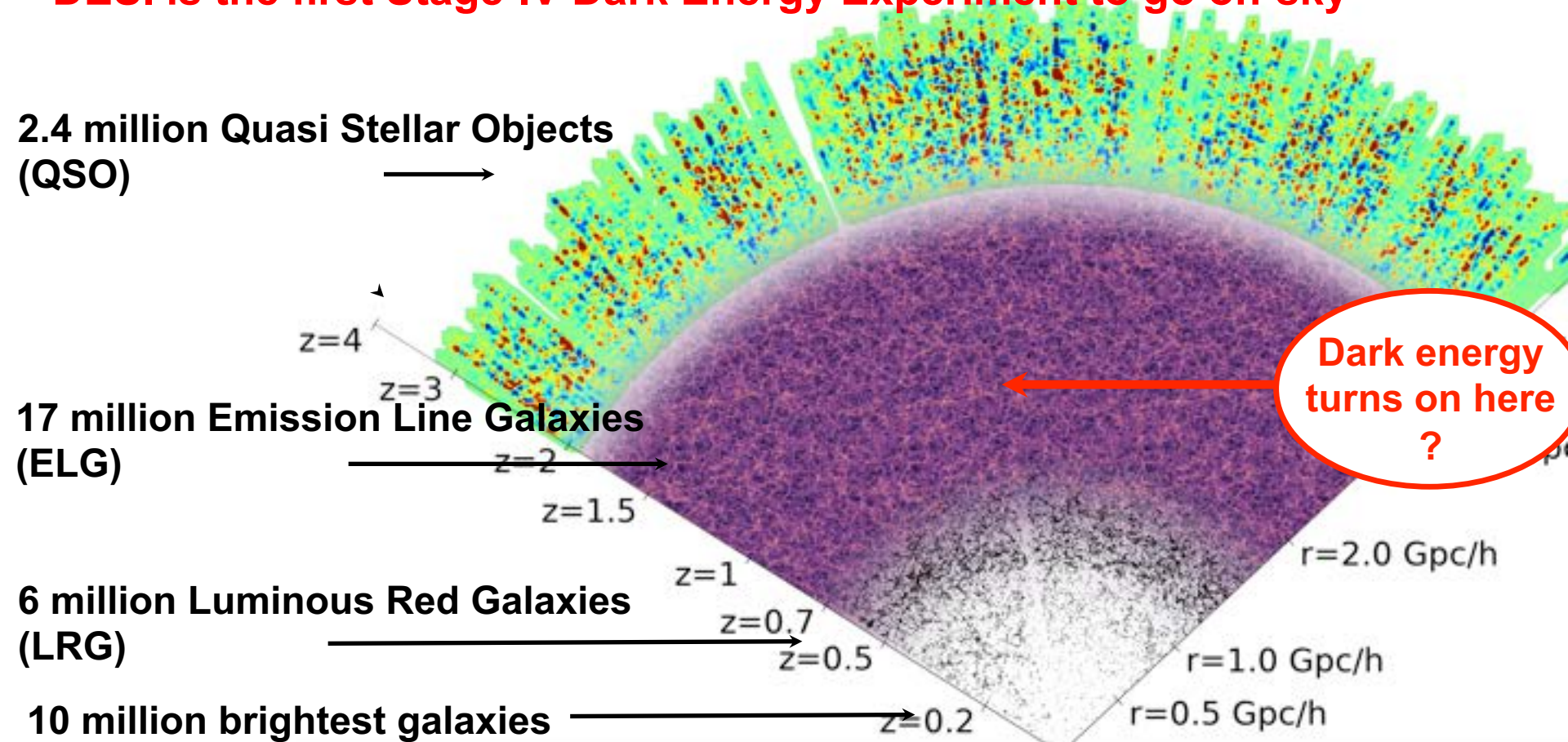
Waterloo CCAT-prime meeting, April 9, 2020

Will Percival

Slide 2

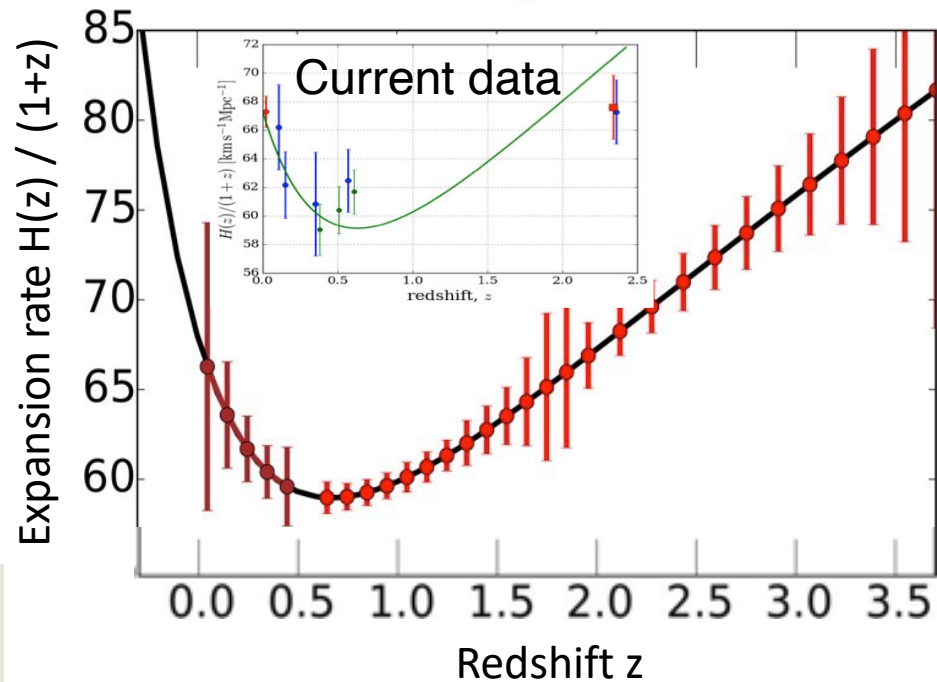
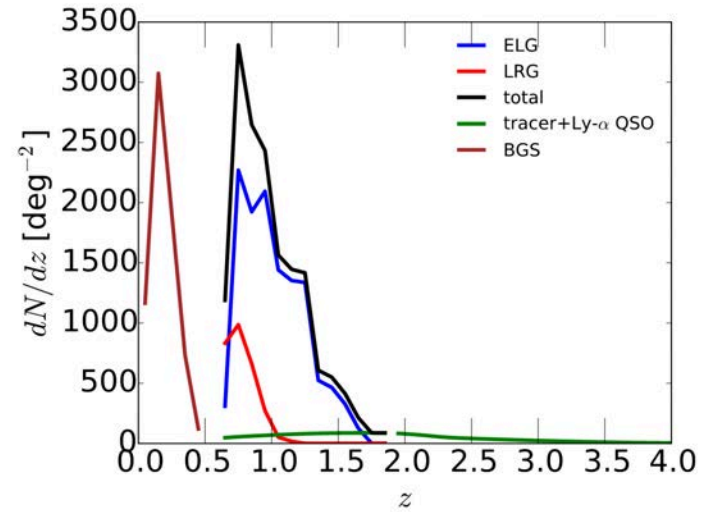
DESI - the largest spectroscopic survey for dark energy

DESI is > 20x faster than SDSS (factor 5 from fibers, 4 from telescope)
DESI is the first Stage-IV Dark Energy Experiment to go on-sky



DESI will collect 35 million spectra

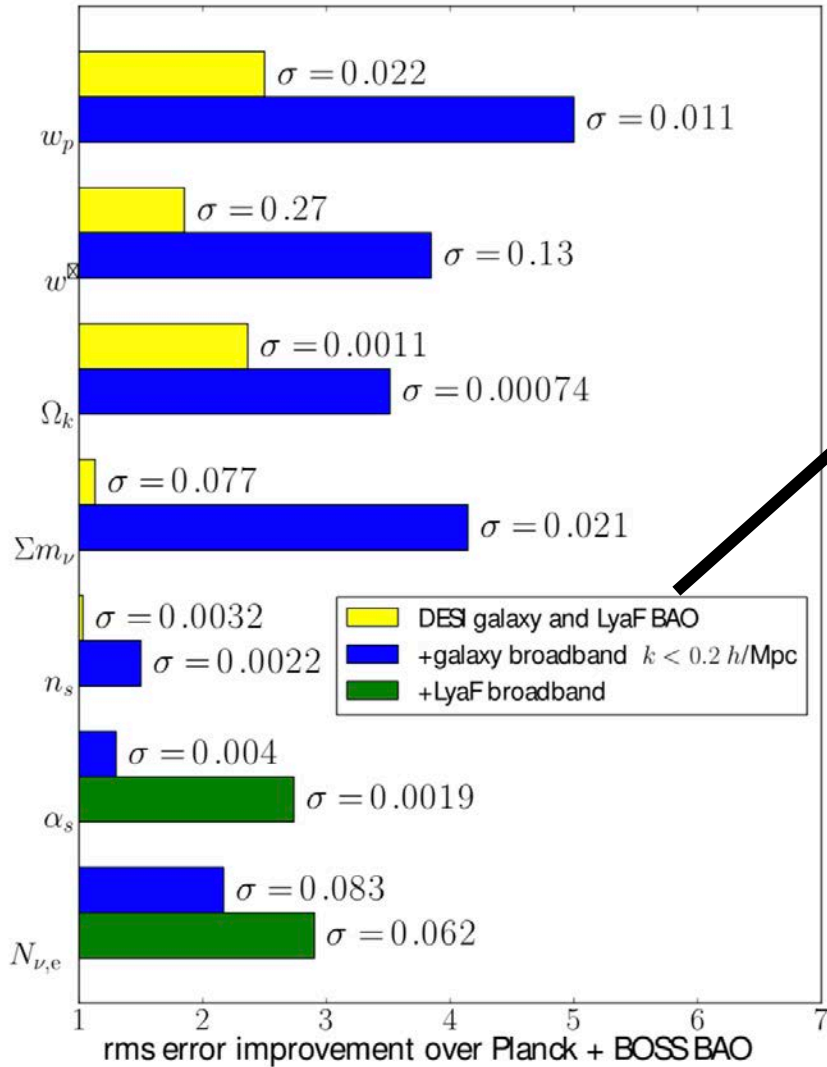
- 5 classes of observation
 - Bright Galaxies
 - Luminous Red galaxies
 - Emission Line Galaxies
 - Quasars as tracers
 - High redshift quasars for Ly- α forest
- Geometrical measurements using Baryon Acoustic Oscillations
- Growth measurements using Redshift Space Distortions
- Lots of other science, including cross-correlations with other surveys (including CCAT-prime)



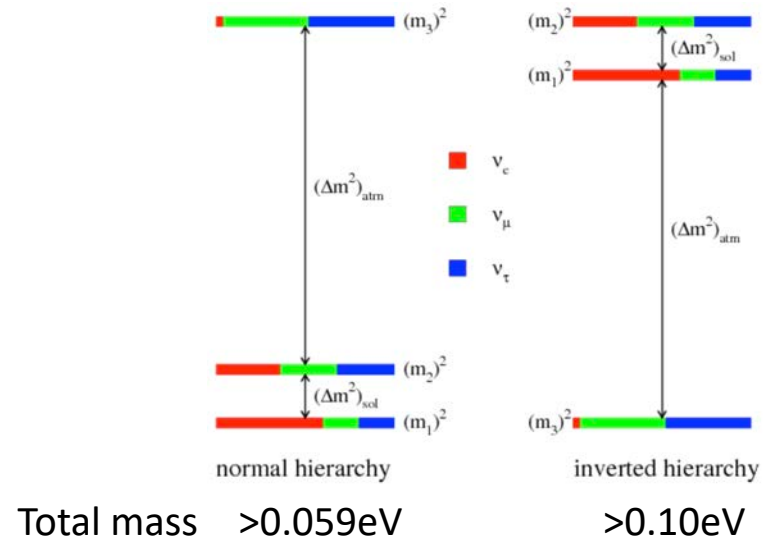
Dark Energy Spectroscopic Instrument

U.S. Department of Energy Office of Science
Lawrence Berkeley National Laboratory

DESI cosmological predictions (arXiv:1611.00036)



Data	$\sigma_{\Sigma m_\nu}$ [eV]	$\sigma_{N_{\nu,eff}}$
Planck	0.56	0.19
Planck + BAO	0.087	0.18
Gal ($k_{max} = 0.1h Mpc^{-1}$)	0.030	0.13
Gal ($k_{max} = 0.2h Mpc^{-1}$)	0.021	0.083
Ly- α forest	0.041	0.11
Ly- α forest + Gal ($k_{max} = 0.2$)	0.020	0.062



Dark Energy Spectroscopic Instrument

U.S. Department of Energy Office of Science
Lawrence Berkeley National Laboratory

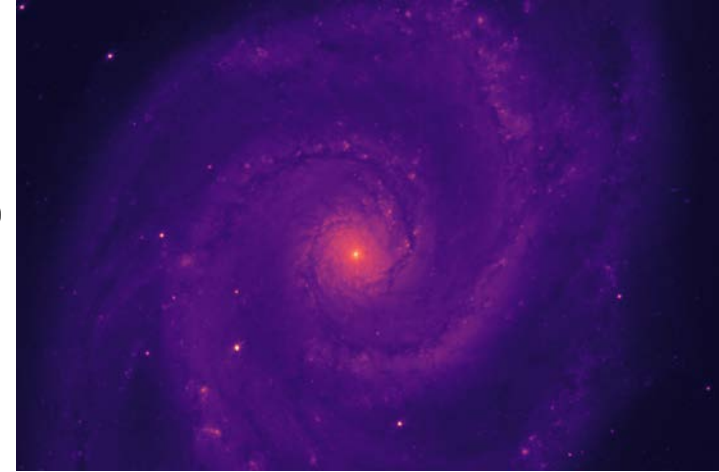
Waterloo CCAT-prime meeting, April 9, 2020

Will Percival

Slide 5

New wide-field optical corrector and focal plane are installed & working at the Mayall Telescope at Kitt Peak

- 6 lenses, largest ~1m in diameter
- First light of corrector images was measured to be 0.7 arcsec (April 1, 2019)
- Whirlpool galaxy as viewed by Commissioning Instrument



Dark Energy Spectroscopic Instrument

U.S. Department of Energy Office of Science
Lawrence Berkeley National Laboratory

Waterloo CCAT-prime meeting, April 9, 2020

Will Percival

Slide 6

Installation and commissioning



Dark Energy Spectroscopic Instrument

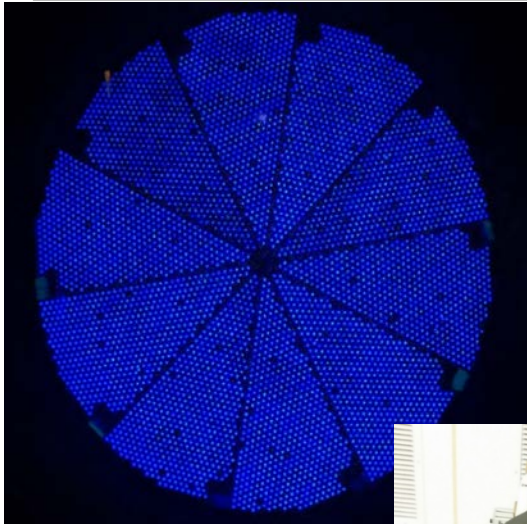
U.S. Department of Energy Office of Science
Lawrence Berkeley National Laboratory

Waterloo CCAT-prime meeting, April 9, 2020

Will Percival

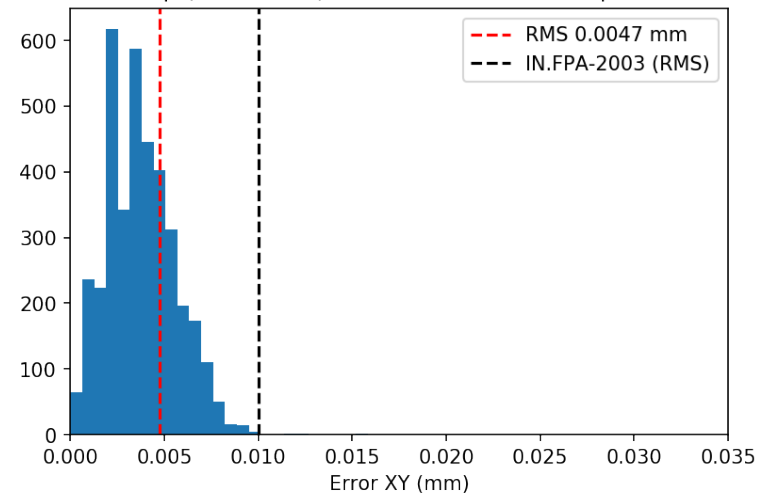
Slide 7

Ten petals constructed



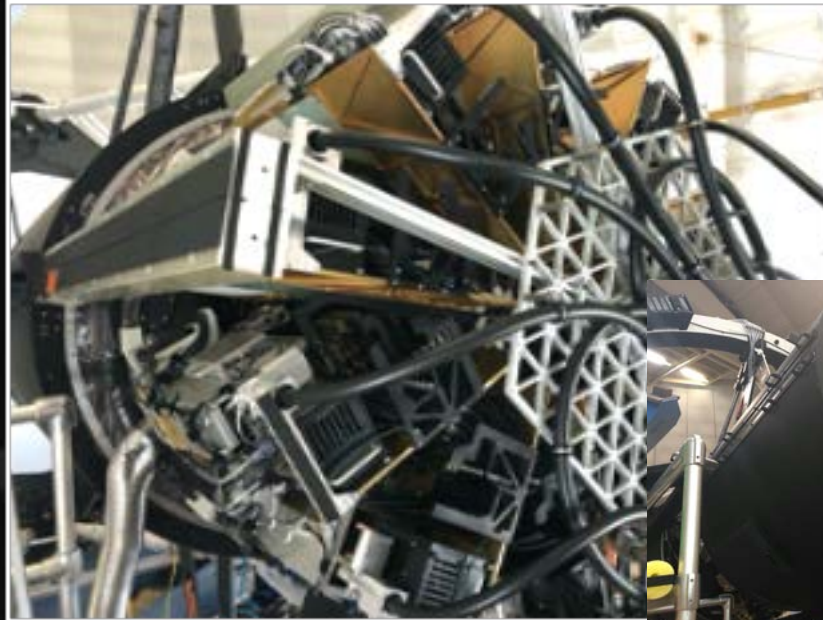
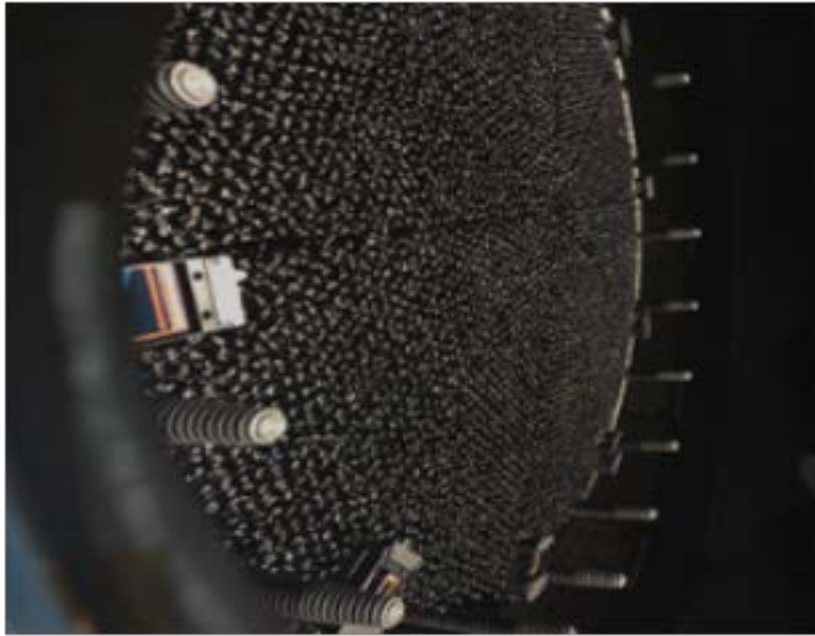
<5 μm rms typical positioning accuracy as-installed

Closed Loop (iteration 2): obsdate 20200306 expid 00053627



Ten petals installed

Installation of the focal plane instrument was completed in August, 2019. The picture below shows the fiber ends of the 5,000 robotic positioners on the focal plane.



Dark Energy Spectroscopic Instrument
U.S. Department of Energy Office of Science
Lawrence Berkeley National Laboratory

Waterloo CCAT-prime meeting, April 9, 2020

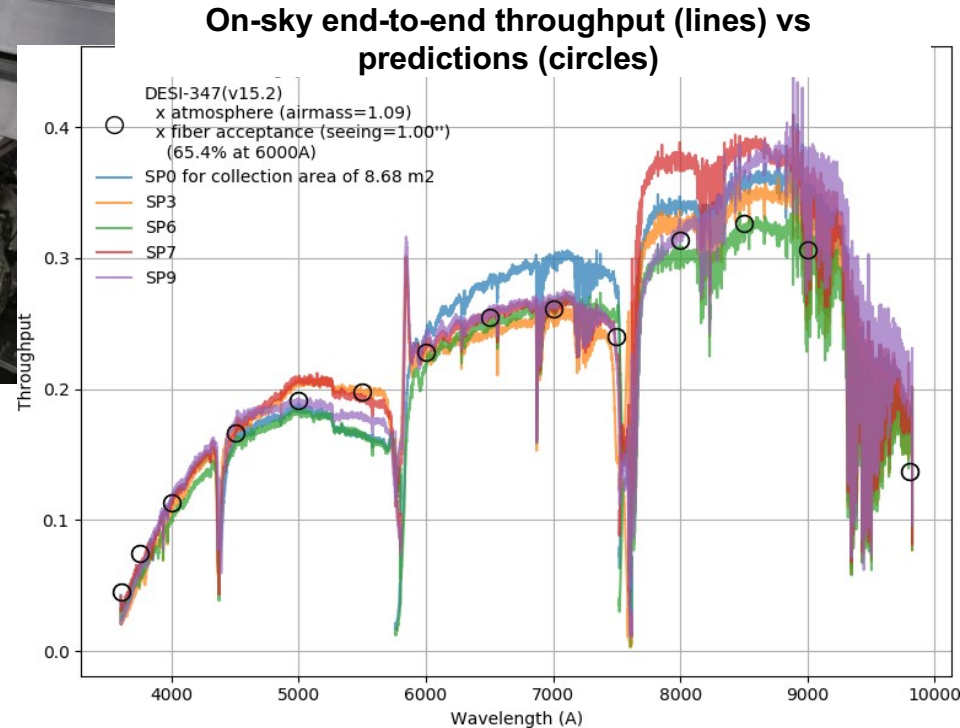
Will Percival

Slide 9

Ten spectrographs installed



10 DESI spectrographs
30 cryostats, 30 CCDs
500 million pixels!!



Dark Energy Spectroscopic Instrument
U.S. Department of Energy Office of Science
Lawrence Berkeley National Laboratory

Waterloo CCAT-prime meeting, April 9, 2020

Will Percival

Slide 10

14,000 sq deg of Imaging data for targeting

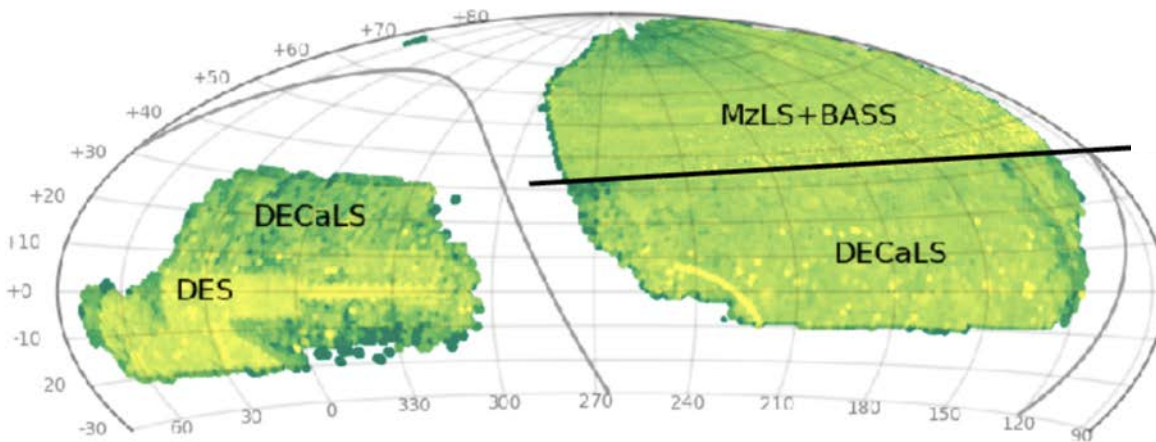
All data public - <http://legacysurvey.org>

Three optical surveys completed
~1.5 mags deeper than SDSS

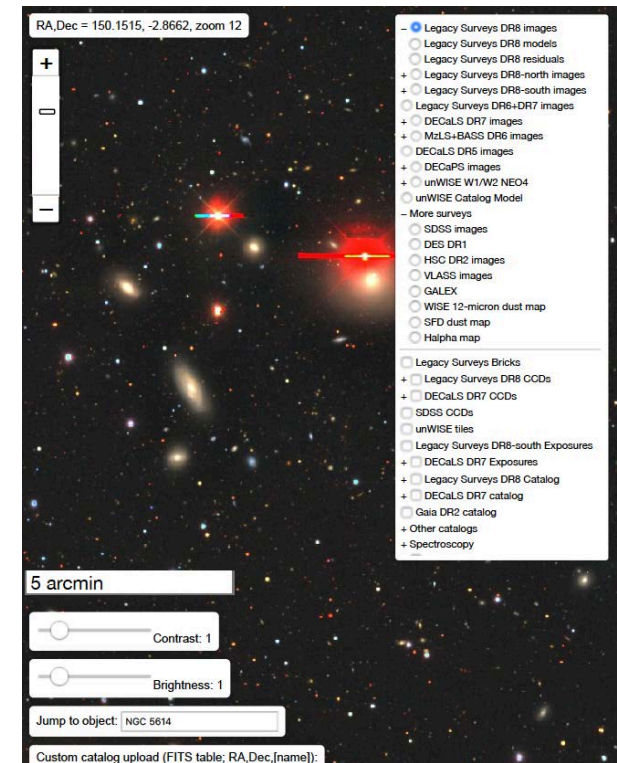
- **North**
(5k deg²) **BASS** gr-bands
- **South**
(9k deg²) **MzLS** z-band
- **DECaLS** grz-bands

One infrared survey completed

- **All Sky WISE** (NASA satellite)
W₁ W₂ bands



- Images combined using Tractor code
- DR8 published July 2019
- Final data release by Q2 2020
- Viewer and other tools available:



Dark Energy Spectroscopic Instrument

U.S. Department of Energy Office of Science
Lawrence Berkeley National Laboratory

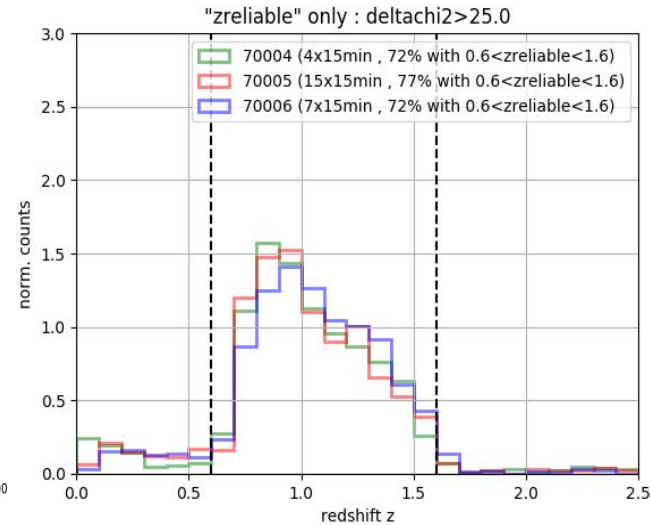
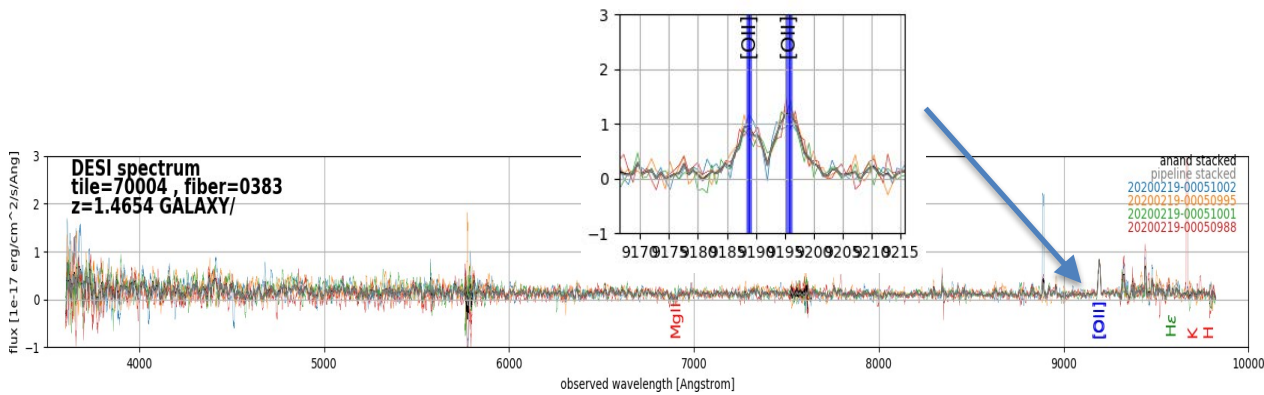
Waterloo CCAT-prime meeting, April 9, 2020

Will Percival

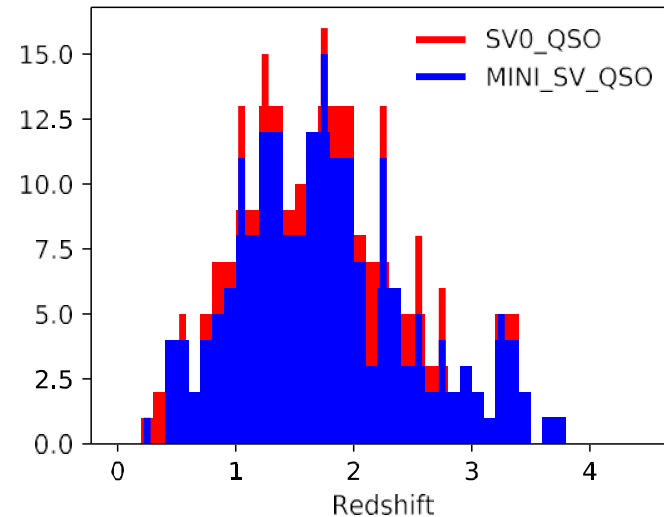
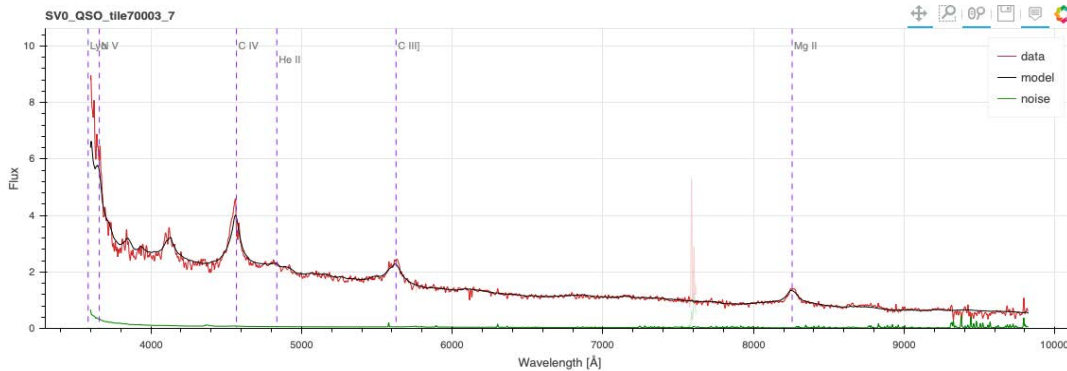
Slide 11

Mini-Survey Validation

- Emission Line Galaxies (ELG)



- Quasi-Stellar Objects (QSO)



Waterloo CCAT-prime meeting, April 9, 2020

Will Percival

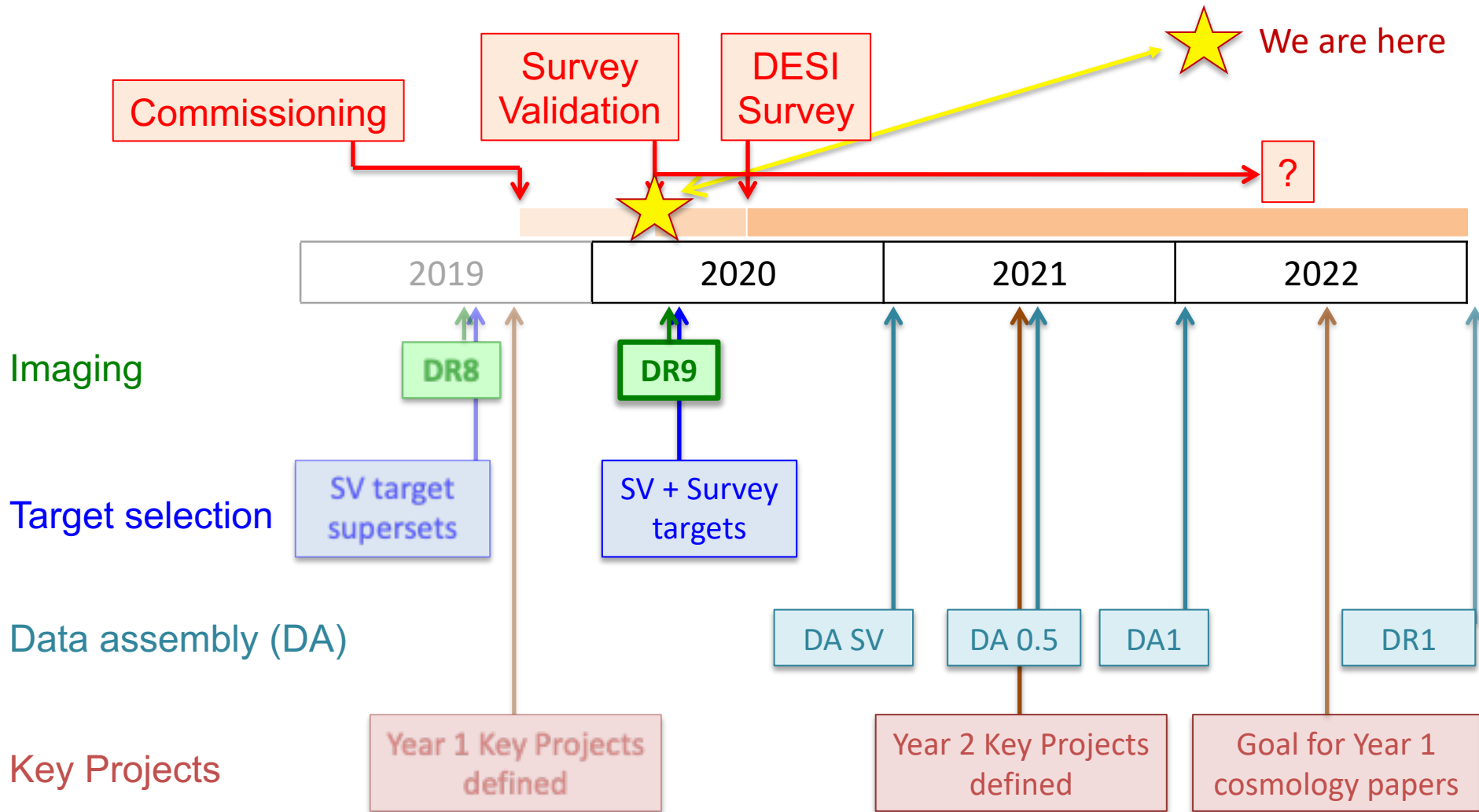
Slide 13

Dark Energy Spectroscopic Instrument

U.S. Department of Energy Office of Science
Lawrence Berkeley National Laboratory



Science can start as soon as safety permits



Dark Energy Spectroscopic Instrument

U.S. Department of Energy Office of Science
Lawrence Berkeley National Laboratory

Waterloo CCAT-prime meeting, April 9, 2020

Will Percival

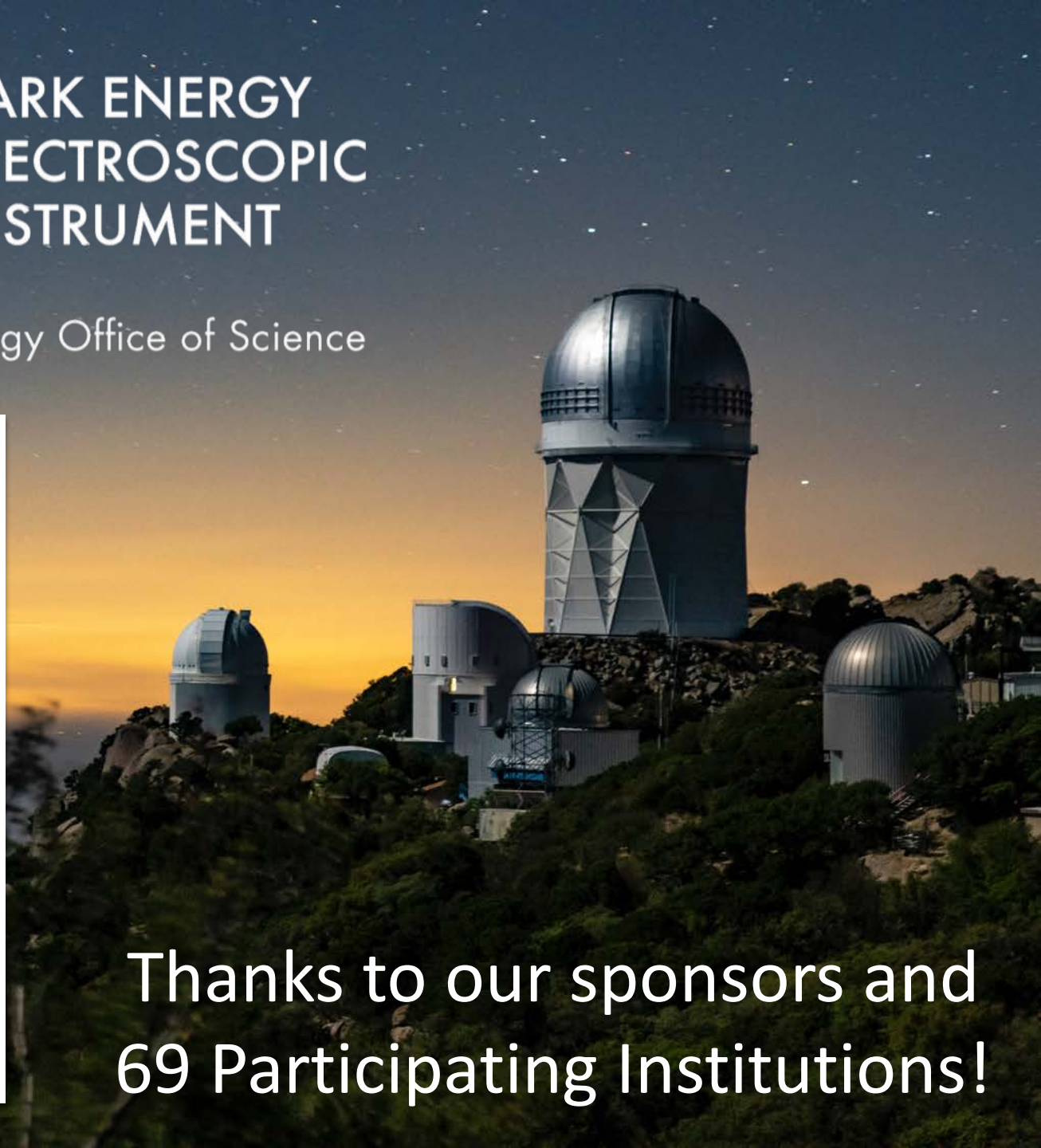


DARK ENERGY SPECTROSCOPIC INSTRUMENT

U.S. Department of Energy Office of Science



Science & Technology
Facilities Council



Thanks to our sponsors and
69 Participating Institutions!